



LAXPC Calibration: A User Perspective

Ranjeev Misra (IUCAA)

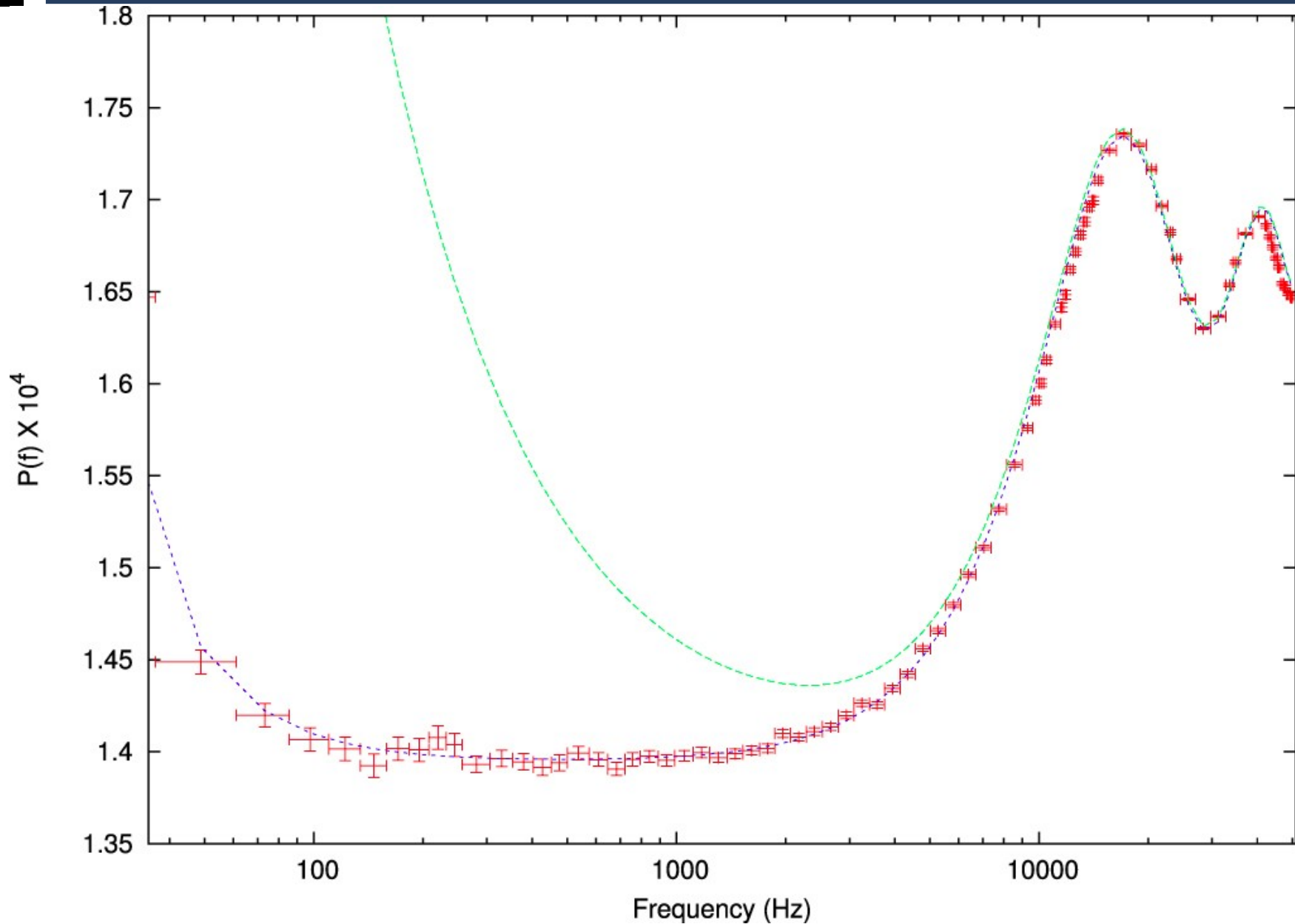


Rapid Timing: VERY GOOD NEWS

- LAXPC rapid timing (< 100 s) is “perfect”
- Except for a coherent narrow feature at 50 Hz for LAXPC 20 (layer 1), which is easily identifiable..

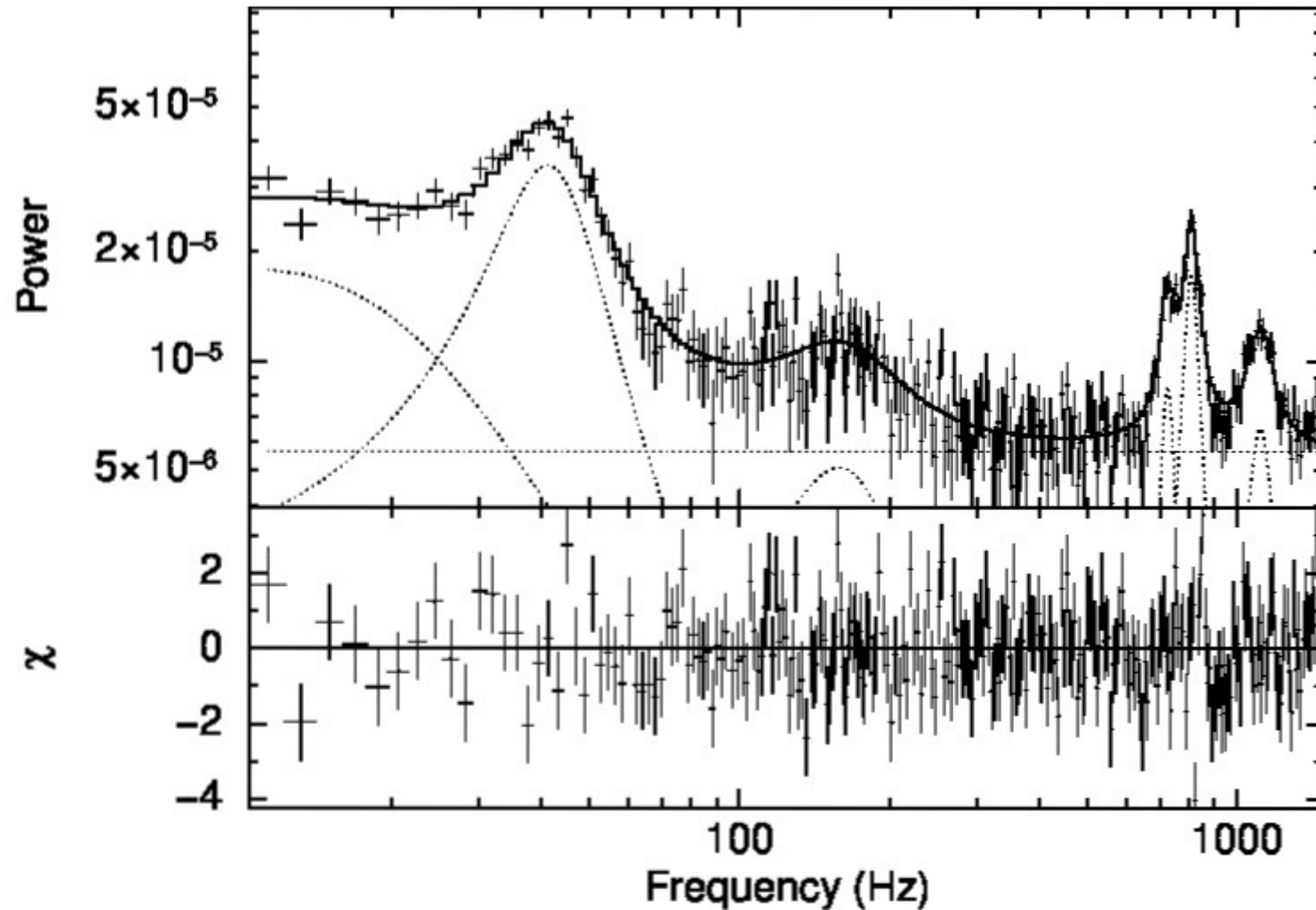


Rapid Timing: VERY GOOD NEWS





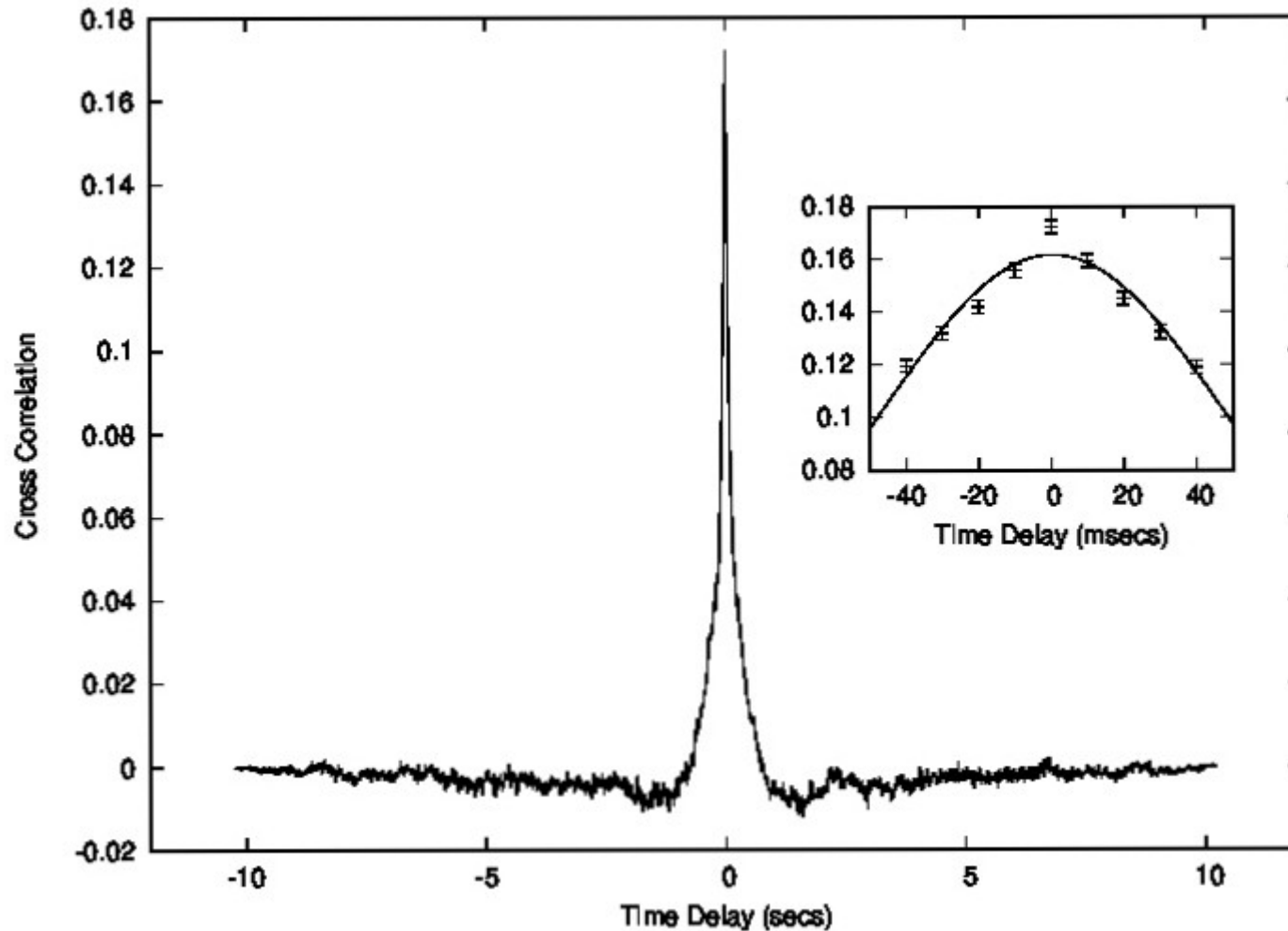
Rapid Timing: VERY GOOD NEWS



CREDIT: Kewal Anand, J. S. Yadav et. al.



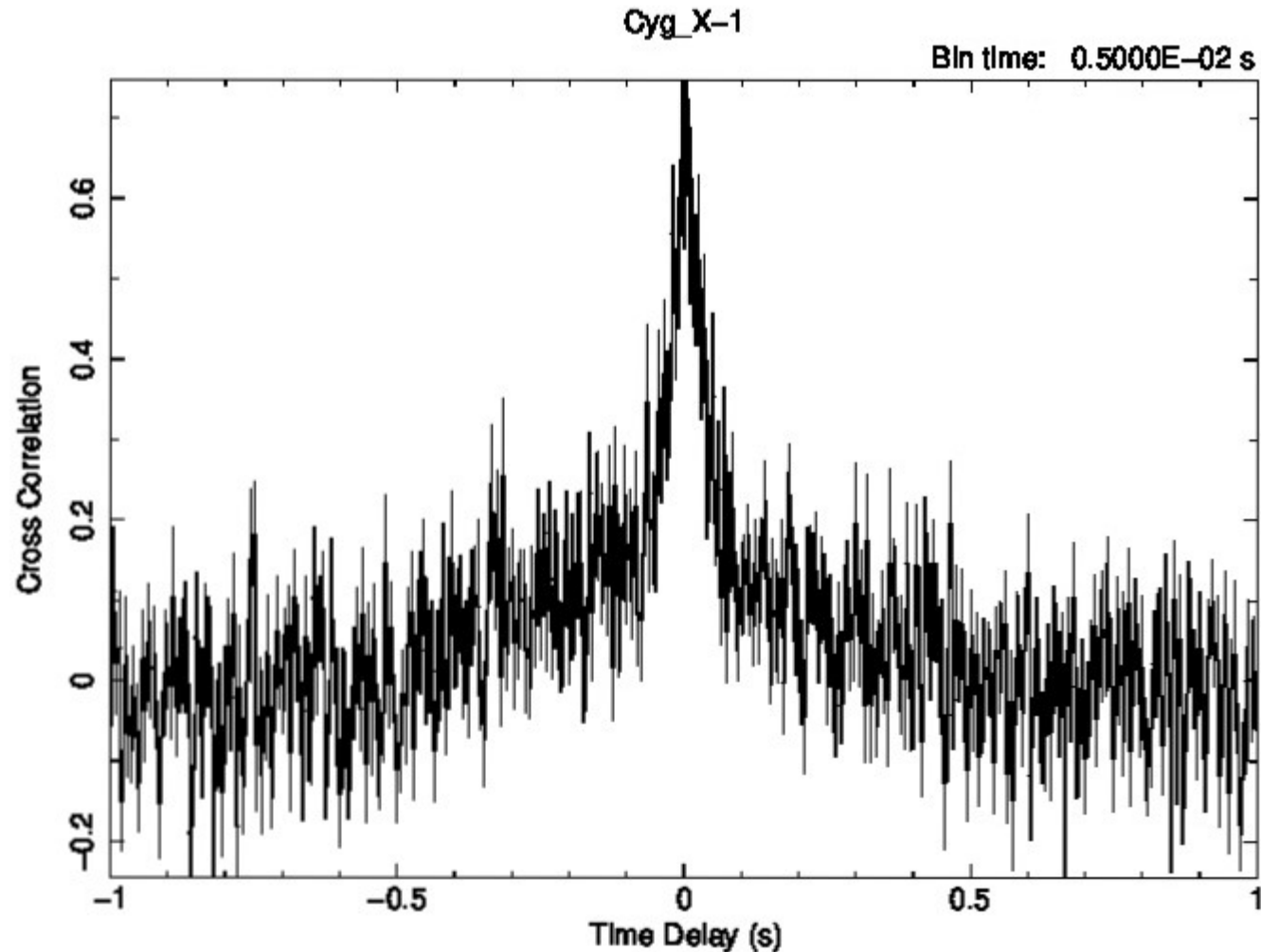
Absolute Timing: VERY GOOD NEWS



NICER and LAXPC...cross correlation



Absolute Timing: *VERY GOOD NEWS*



HMXT ME and LAXPC...cross correlation 10-20 keV
CREDIT: Gitika Mall et. al.



Rapid Timing: VERY GOOD NEWS

- LAXPC rapid timing (< 100 s) is “perfect”
- Except for a coherent narrow feature at 50 Hz for LAXPC 20 (layer 1), which is easily identifiable..
- **THANK YOU TO THE INSTRUMENT TEAM FROM THE USER COMMUNITY!!!!**

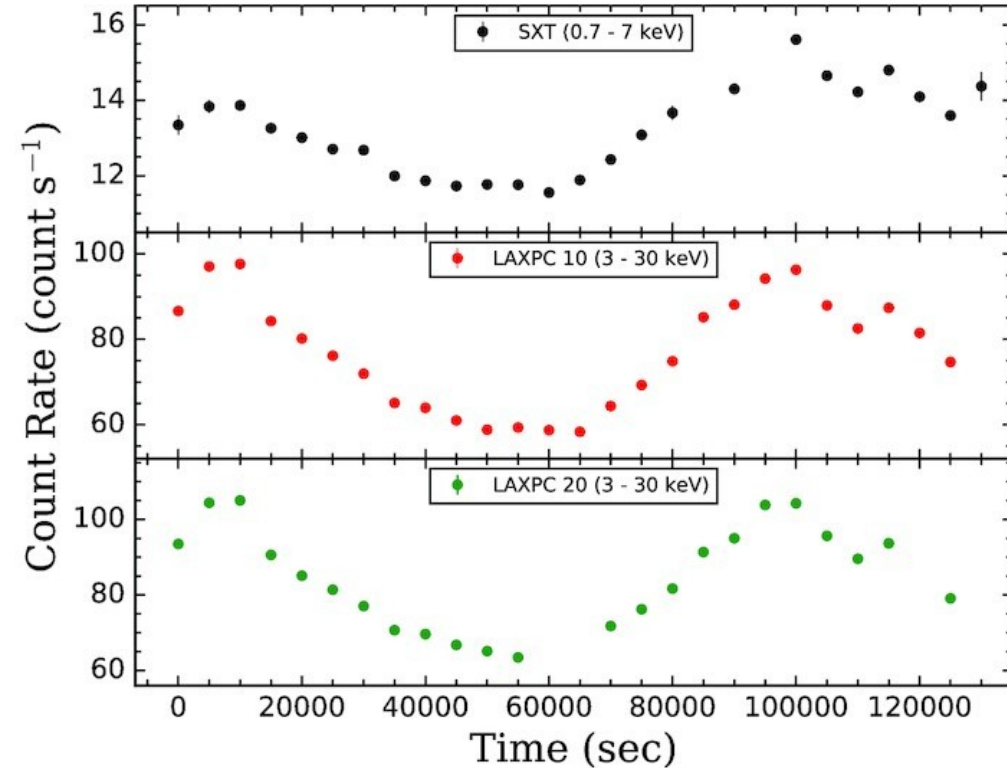
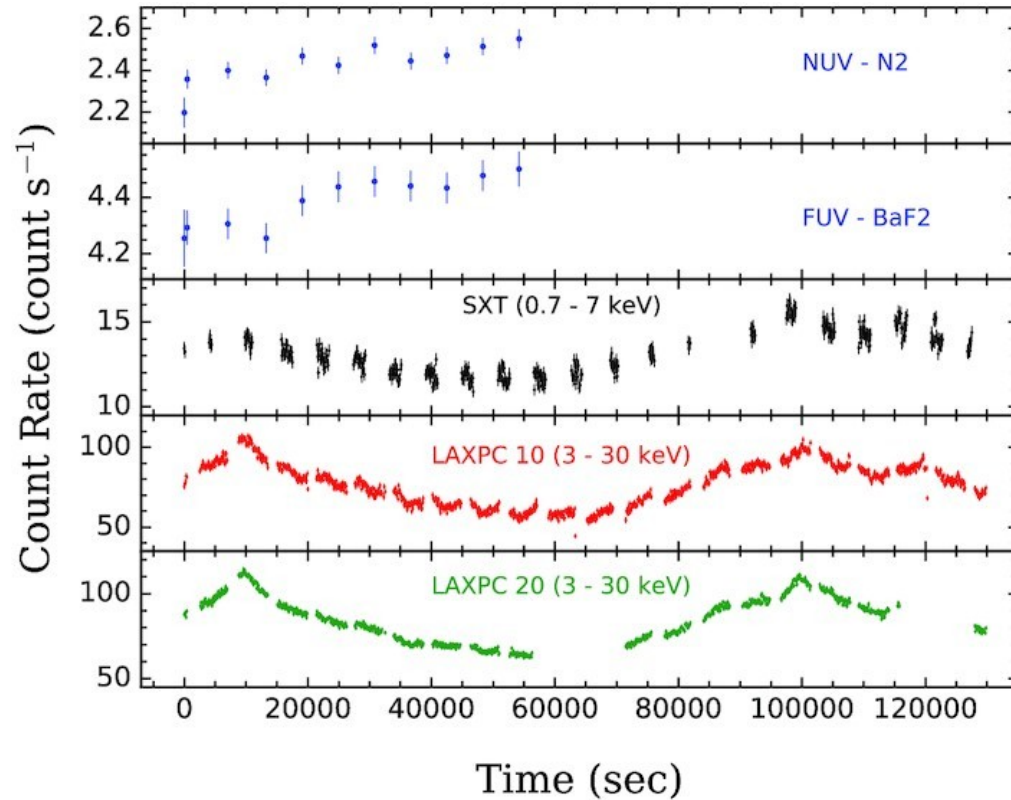


Long term studies: GOOD NEWS

- For bright highly variable sources..no issues..
- For faint sources ..background variability needs to be considered..
 - ➔ New background model (2021)..very helpful..
- SXT lightcurve can be used to validate results



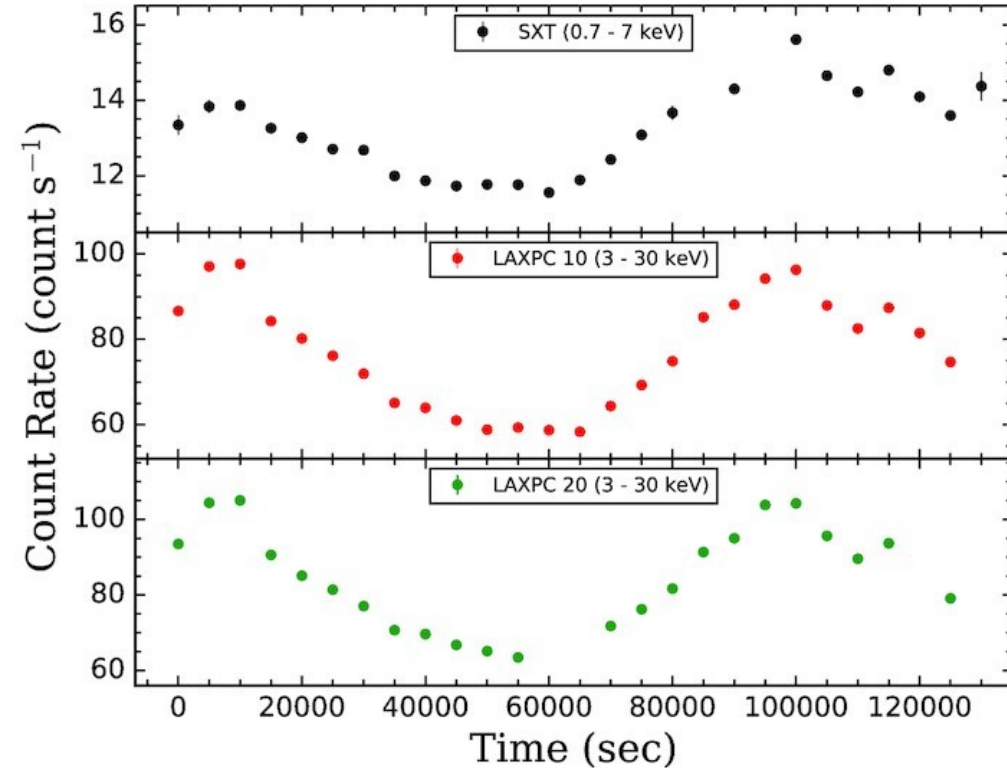
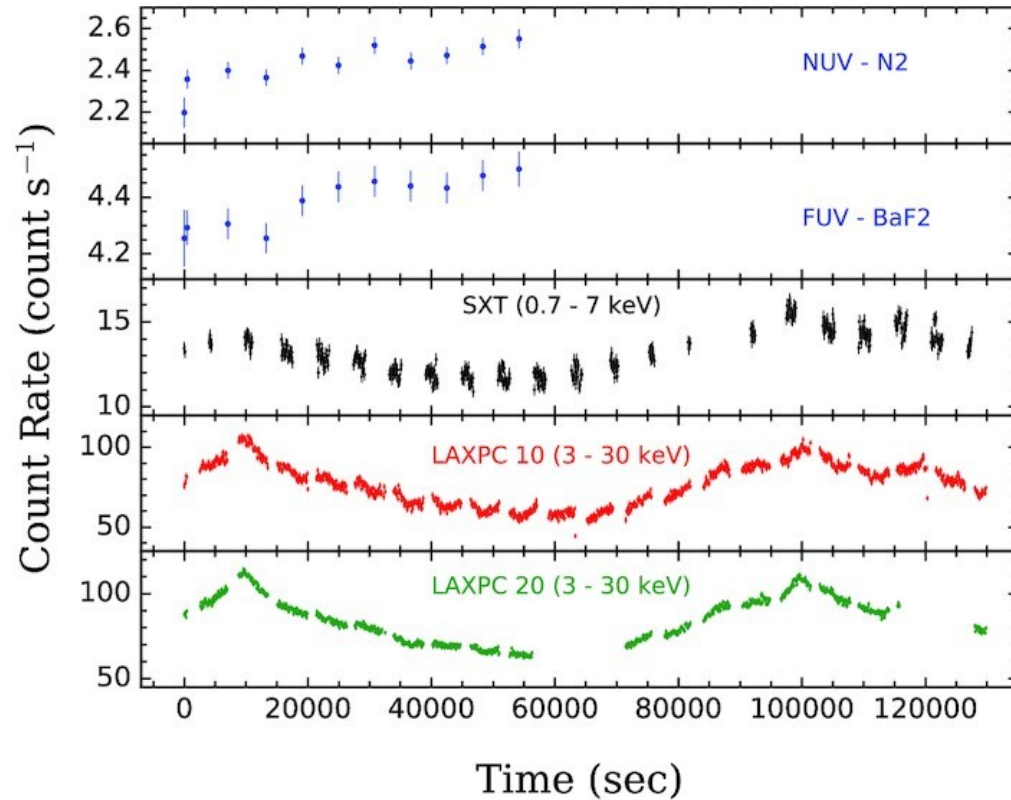
Long term studies: *GOOD NEWS*



Shah et. al...Data from a blazar



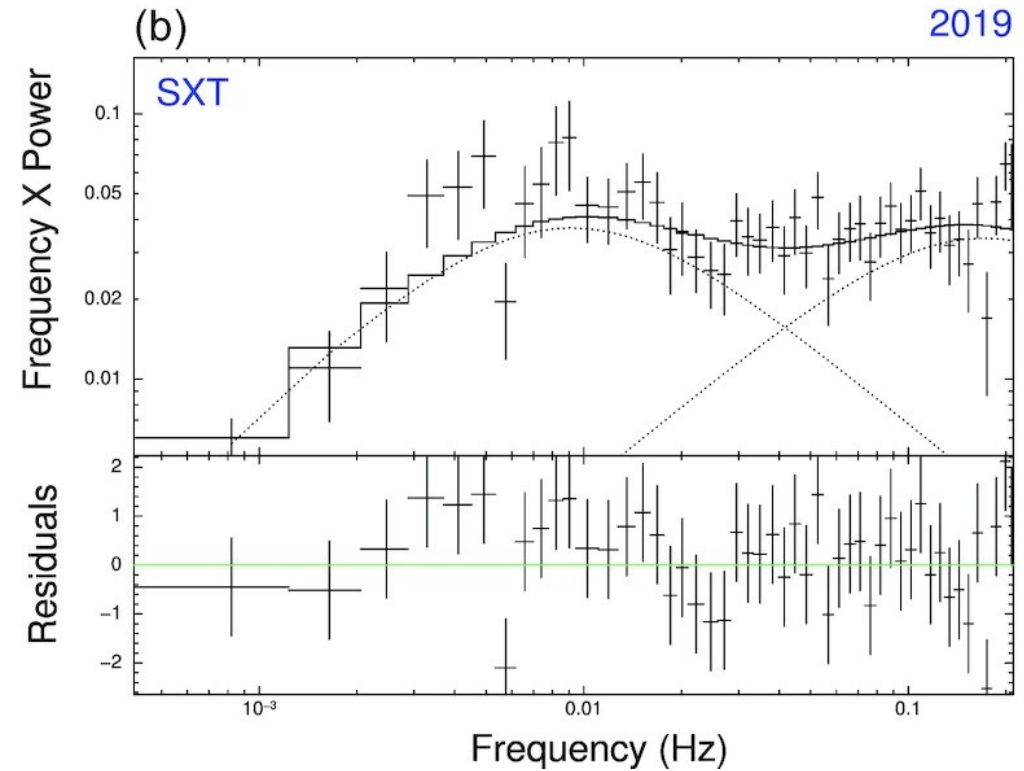
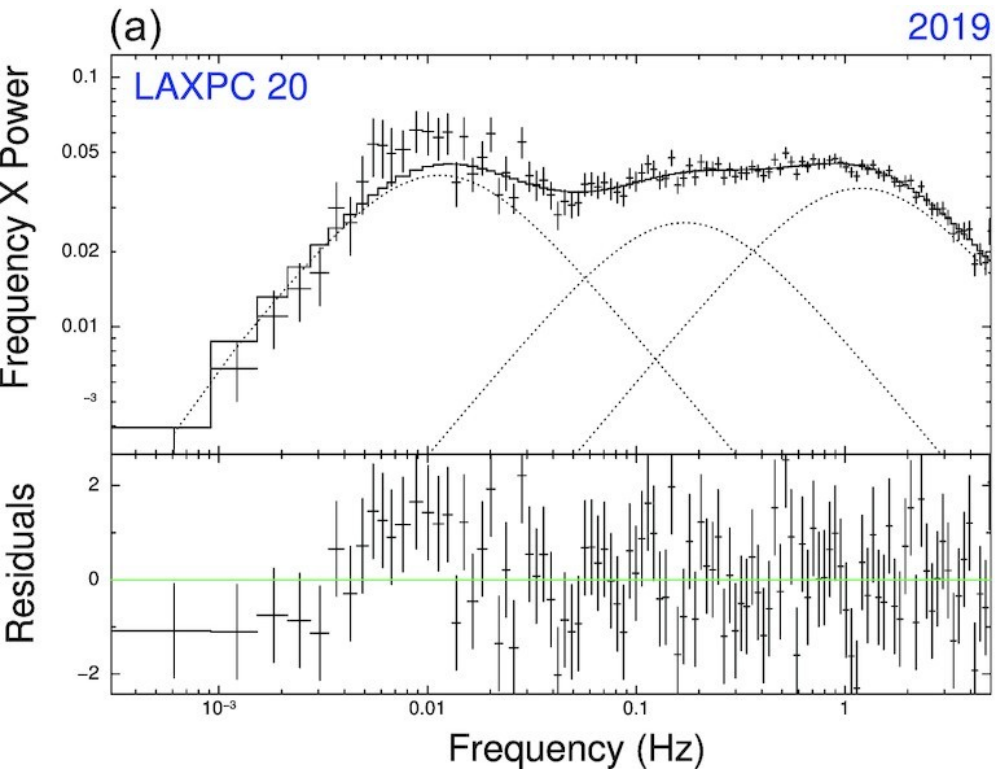
Long term studies: *GOOD NEWS*



Shah et. al...Data from a blazar



Long term studies: *GOOD NEWS*



Husain et. al...Data from an X-ray binary



Faint Source Monitoring..

- Poisson limit: 1 milli-crab source flux/spectral index can be constrained for a 5 ksec observation.
- At present, the systematic error on background is of the same order as the Poisson error for 5 ksec.
- Thus, monitoring of a source with with few millicrab flux is feasible and worth doing.
- The flux limit can be decreased by better handle on background systematics

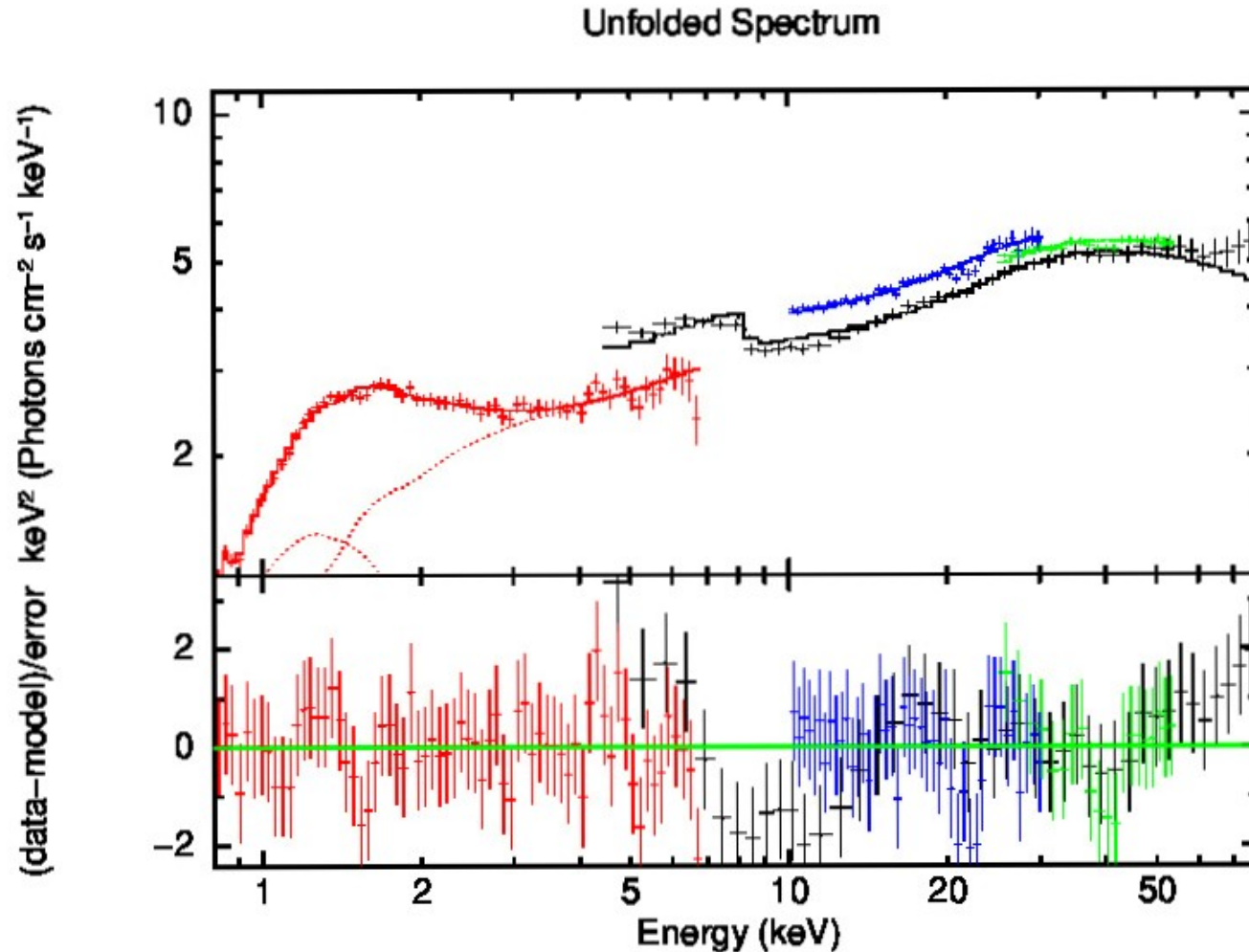


Spectral Studies: GOOD NEWS

- Systematics have to be included for spectra (3 % or 2 % or...)
- Systematics have to be included for background (3 % or 2 % or..)
- For most X-ray binary and AGN spectral continuum studies provides good results..
 - ➔ Determined by SXT spectra..



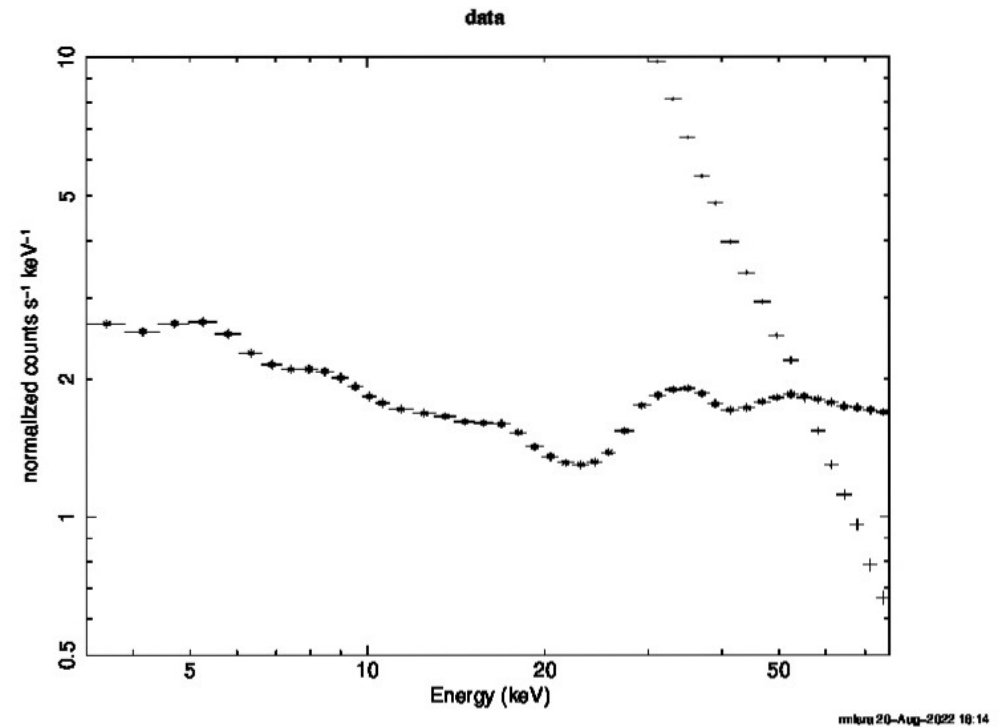
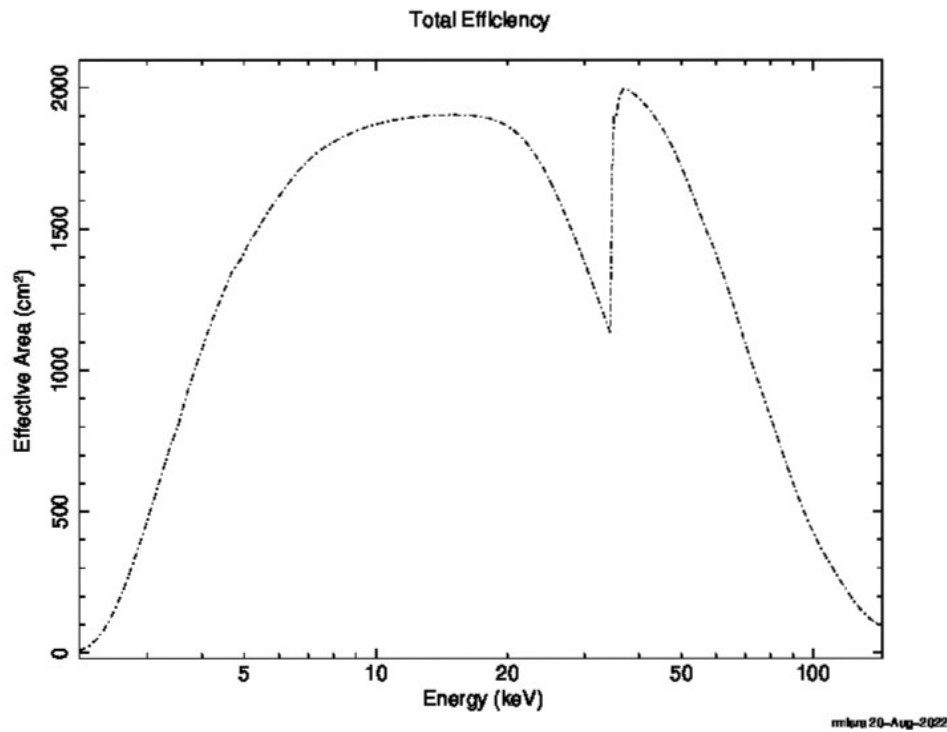
Spectral studies: Good News..



LAXPC (BLACK), SXT (RED), HXMT ME (BLUE), HXMT HE (GREEN)
PRELIMINARY RESULT.. Figure by Gitika Mall..



Spectral studies: Not so good news..





Spectral studies: Not so good news..

- Study of pulsar cyclotron absorption line at around 30 keV is compromised..
- LAXPC is the only instrument that can in principle do phase resolved cyclotron line analysis of rapid pulsars...

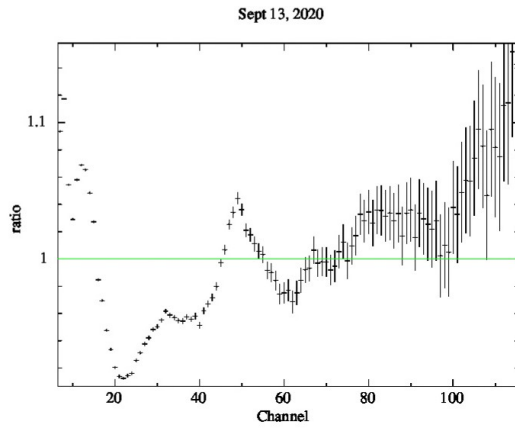


An idea and suggestions....

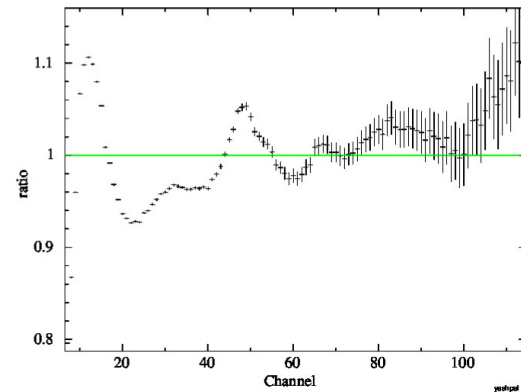
- Possibility of using Crab observations that are a few days close to the source observation, to empirically improve the spectral fitting..



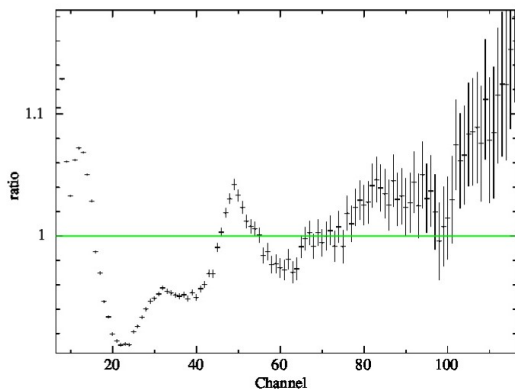
Crab Analysis..Preliminary: Credit: Yashpal Bhulla



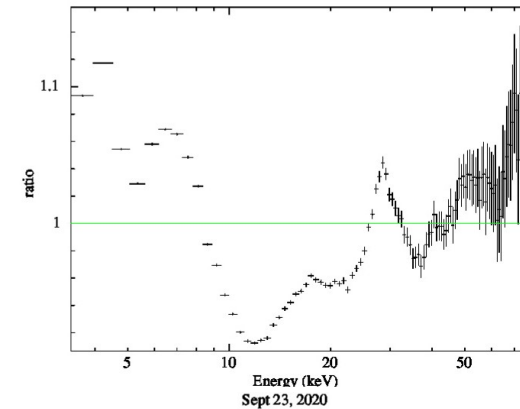
Sept 23, 2020



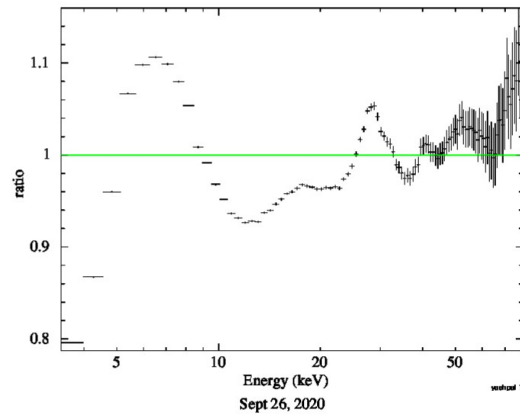
Sept 26, 2020



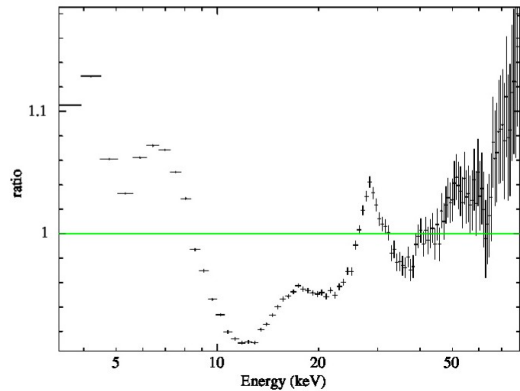
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Sept 23, 2020

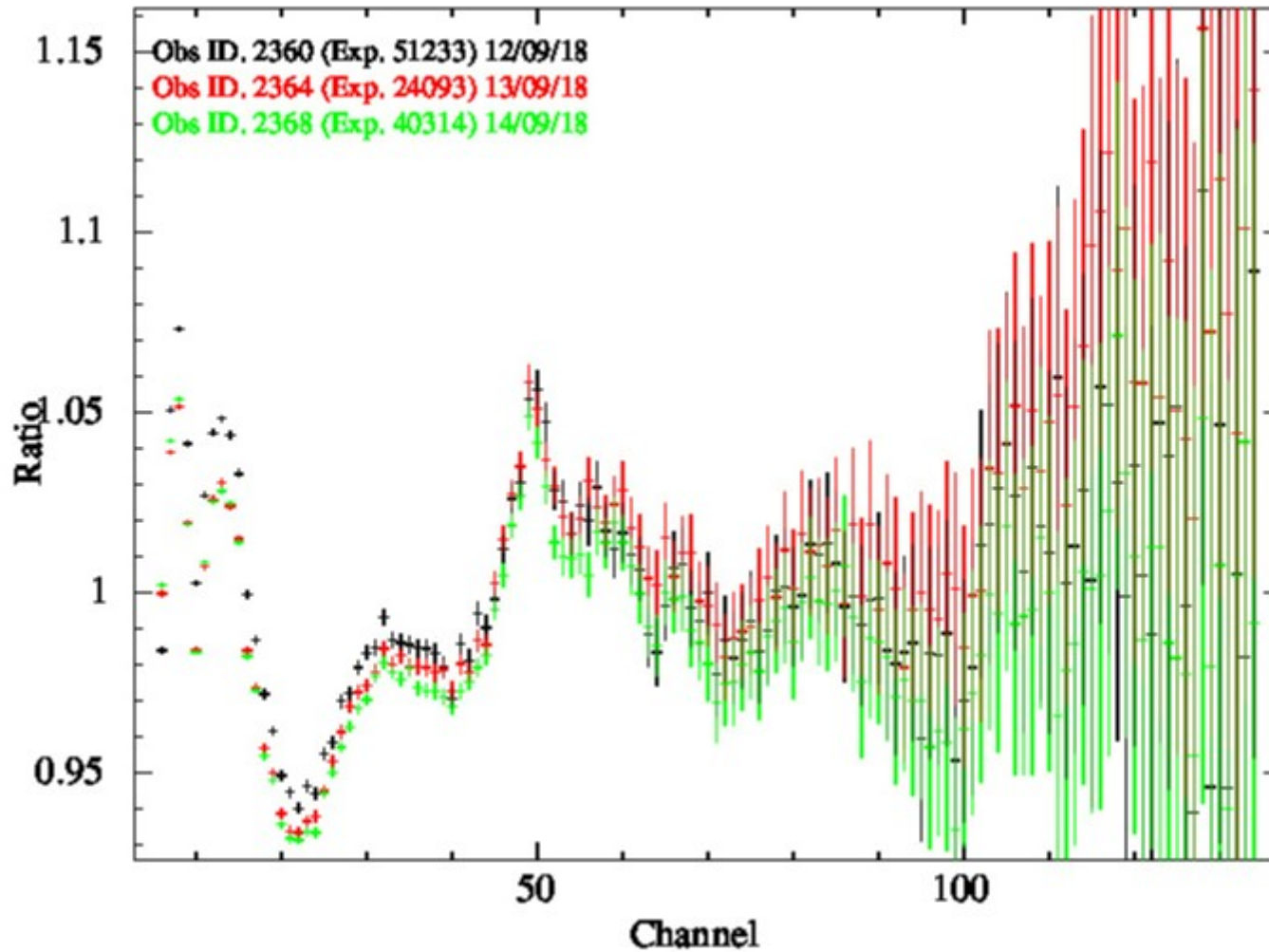


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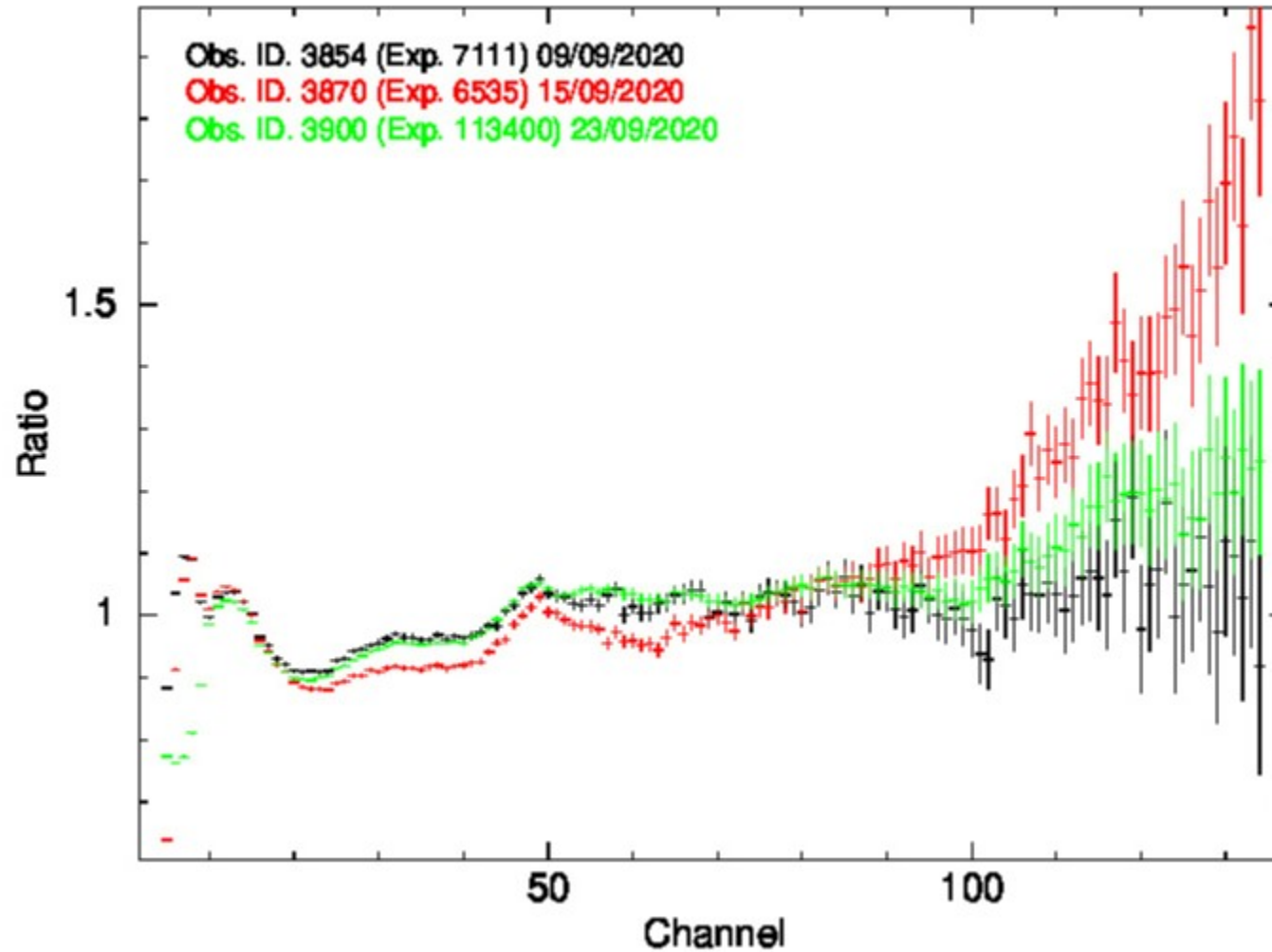


Crab Analysis..Preliminary: Credit: Yashpal Bhulla





Crab Analysis..Preliminary: Credit: Yashpal Bhulla





An idea...the plan..

- Observe a pulsar to study phase resolved cyclotron absorption line..
- Follow it up immediately with a Crab observation...
- Correct the source spectra based on Crab residuals...



An idea....Verification....

- Observe any bright source that Nustar is observing..
- Follow it up immediately with a Crab observation...
- Check if the source spectra corrected based on Crab residuals, matches with the Nustar one...



An idea and suggestions....

- Possibility of using Crab observations that are a few days close to the source observation, to empirically improve the spectral fitting..



Suggestions and an idea....

- Possibility of using Crab observations that are a few days close to the source observation, to empirically improve the spectral fitting..
- Energy dependent systematic errors that can be included in the spectral file.



Suggestions and an idea....

- Possibility of using Crab observations that are a few days close to the source observation, to empirically improve the spectral fitting..
- Energy dependent systematic errors that can be included in the spectral file.
- Exposure time dependent systematic error for background (energy dependent?)



THANK YOU



Faint source detection:

